ExoPlanet News An Electronic Newsletter

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1 EDITORIAL

1 Editorial

Welcome to Edition 193 of ExoPlanet News!

This month, as usual, we bring you abstracts of scientific papers, job ads, conference announcements, and an overview of exoplanet-related articles on astro-ph. Thanks a lot to all of you who contributed to this issue of the newsletter!

For next month, we look forward to continuing receiving your submissions of paper abstracts, job ads, or meeting announcements. Special announcements are also welcome. As always, we would also be happy to receive feedback concerning the newsletter. The LAT_EX template (v2.0) for submitting contributions, as well as all previous editions of ExoPlanet News, can be found on the ExoPlanet News webpage (https://nccr-planets.ch/exoplanetnews/).

The next issue will appear on Tuesday, August 12th (with a submission deadline ending on Sun August 10th, 2025 CEST).

Thanks again for your support, and best wishes from the editorial team.

Leander Schlarmann Haiyang Wang Jeanne Davoult Timm-Emanuel Riesen



Univ. of Bern, Univ. of Geneva, ETH Zürich, Univ. of Zürich, EPF Lausanne The National Centers of Competence in Research (NCCR) are a research instrument of the Swiss National Science Foundation.

2 ABSTRACTS OF REFEREED PAPERS

2 Abstracts of refereed papers

Three hot Jupiters transiting K-dwarfs with significant heavy element masses

Y.G.C. Frensch¹, F. Bouchy¹, G. Lo Curto², S. Ulmer-Moll³, S.G. Sousa⁴, N.C. Santos^{4,5}, K.G. Stassun⁶, C.N. Watkins⁷, H. Chakraborty¹, et al. (a complete list of authors can be found on the publication)

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A&A, in press (arXiv:2506.04923)

We report the confirmation and characterization of three giants transiting mid-K dwarfs. Within the TOI-2969 system, a giant planet of $1.16 \pm 0.04 M_{Jup}$ with a radius of $1.10 \pm 0.08 R_{Jup}$ orbits its K3V host in 1.82 days. The TOI-2989 system contains a $3.0 \pm 0.2 M_{Jup}$ giant with a radius of $1.12 \pm 0.05 R_{Jup}$, which orbits its K4V host in 3.12 days. The K4V star TOI-5300 hosts a giant of $0.6 \pm 0.1 M_{Jup}$ with a radius of $0.88 \pm 0.08 R_{Jup}$ and an orbital period of 2.3 days. The equilibrium temperatures of the companions range from 1001 to 1186 K, which classifies them as hot Jupiters. However, they do not exhibit radius inflation. The estimated heavy element masses in their interiors, inferred from the mass, radius, and evolutionary models, are $90 \pm 30 M_{\oplus}$, $114 \pm 30 M_{\oplus}$, and $84 \pm 21 M_{\oplus}$, respectively. These heavy element masses are significantly higher than most reported heavy elements for K-dwarf hot Jupiters.

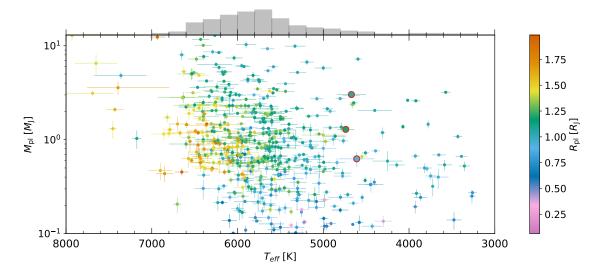


Figure 1: Overview of the presented companions (encircled in red) compared to known planets from the PlanetS catalog. The data are color coded by planetary radii. The histogram at the top shows the relative occurrence of the transiting gas giants with masses ranging from 0.1 to 13 $M_{\rm Jup}$. The low-mass star regime remains relatively poorly populated; our three mass characterizations contribute to this population.

Download/Website: https://arxiv.org/abs/2506.04923 Contact: yolanda.frensch@unige.ch

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2 ABSTRACTS OF REFEREED PAPERS

Feasibility of interferometric observations and characterization of planet-induced structures at sub au to au scales in protoplanetary disks

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Astronomy & Astrophysics, published (2025A&A...698A..67H)

Interferometric observations of protoplanetary disks by VLTI and ALMA have greatly improved our understanding of the detailed structure of these planetary birthplaces. These observations have revealed a variety of large-scale disk substructures, including rings, gaps, and spirals, spanning tens to hundreds of au, supporting the predictions of planet formation models. Recent instruments, such as MATISSE at the VLTI, allow one to resolve and investigate the inner few au of protoplanetary disks in nearby star formation regions, shedding light on the traces of planet formation and evolution at these small scales. The aim of this work is to assess the feasibility of interferometric observations of small-scale planet-induced substructures in protoplanetary disks in nearby star-forming regions. We aim to characterize these substructures in multi-wavelength and multi-epoch observations and subsequently differentiate between simulation parameters. On the basis of 3D hydrodynamic simulations of embedded planetary companions and subsequent 3D Monte Carlo radiative transfer simulations, we calculated and analyzed interferometric observables, assuming observations with the VLTI in the K, L, M, and N bands. The hydrodynamic simulations exhibit mass-dependent planet-induced density waves that create observable substructures, most notably for the considered case of a 300 M_{\oplus} planet. These substructures share similarities with observed large-scale structures and feature a prominent accretion region around the embedded planet. The visibilities show a detectable variability for multi-epoch VLTI/GRAVITY and VLTI/MATISSE observations, caused by the orbital motion of the planet, that are distinguishable from other sources of variability due to their unique combination of timescale and amplitude. Additionally, the non-uniform change of the visibilities at different baselines can be used to identify asymmetric structures. Furthermore, we show that multi-wavelength observations provide an approach to identify the fainter substructures and the signal of the accretion region.

Download/Website: https://arxiv.org/abs/2506.17024 Contact: lhildebrandt@astrophysik.uni-kiel.de

Saturns, but not super-Jupiters, occur more frequently in the presence of inner super-Earths

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² UFR Physique, Université Paris Cité, 4 Rue Elsa Morante, F-75013 Paris, France

The Astrophysical Journal, in press (arXiv:2506.21204)

Studies from recent years have reached different conclusions regarding how frequently super-Earths are accompanied by long period giant planets and vice versa. This relation has been predicted to be mass dependent by planet formation models. We investigate that as the origin of the discrepancy using a radial velocity sample: the California Legacy Survey. We perform detection completeness corrections in order to discard detection bias as a possible explanation to our results. After bias corrections, we find that cold Jupiters are $5.65^{+1.08}_{-2.57}$ times more massive when not in company of an inner super-Earth, while super-Earths are not significantly more massive while in company of an outer giant planet. We also report an occurrence enhancement for Saturns (median projected mass of $0.6M_J$) while in presence of a super-Earth by a factor of ~4, and for super-Earths in presence of Saturns by the same factor. This positive correlation disappears for super-Jupiters (median projected mass of $3.1M_J$). These results show that while cold Jupiters are generally accompanied by inner super-Earths, this does not hold for the largest giant planets, such as those that will be discovered by Gaia, which will likely not be accompanied by transiting planets. The mass dependence, in combination with different detection limits of different surveys, may explain the discrepancies concerning occurrence relations between cold Jupiters and super-Earths.

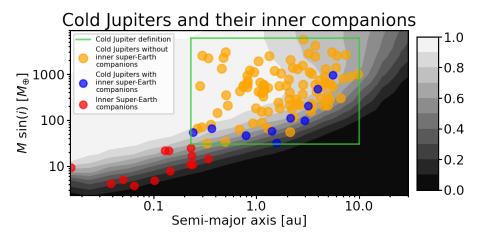


Figure 2: Cold Jupiters and their inner super-Earth companions. The contours correspond to the detection completeness of systems that host at least one cold Jupiter.

Download/Website: https://arxiv.org/abs/2506.21204 Contact: etienne.lefevre.forjan@gmail.com

3 EXOPLANET ARCHIVES

3 Exoplanet Archives

June 2025 Updates at the NASA Exoplanet Archive

The NASA Exoplanet Archive team

Caltech/IPAC-NASA Exoplanet Science Institute, MC 100-22 Pasadena CA 91125

Pasadena CA USA, July 8, 2025

Note: Unless otherwise noted, all planetary and stellar data mentioned in the news are in the Planetary Systems Table, which provides a single location for all self-consistent planetary solutions, and its companion table the Planetary Systems Composite Parameters, which offers a more complete table of parameters combined from multiple references and calculations. Links to other tables and System Overview pages are embedded in the news text.

June 26, 2025

Five Planets and Eight Spectra

This week's release includes Kepler-725 c, a planet of 10 Earth masses located in its Sun-like star's optimistic habitable zone. Check out the media release and Nature paper for details. The other new planets are TOI-1117 b, c, & d and TOI-2407 b.

There are also eight new spectra for TRAPPIST-1 c, WASP-96 b, and the nearby sub-stellar object WISE J1049 b (a.k.a. Luhman 16 b). Access these in the Atmospheric Spectroscopy Table.

June 12, 2025

A Giant Planet Orbiting a Tiny Star

One of this week's four planets is TOI-6894 b, a sub-Saturn-mass planet orbiting a low-mass red dwarf-—the smallest known star to host a transiting giant planet. This discovery challenges some current planetary formation models and theories. Check out the University College London media release and the discovery paper.

The other new planets are TOI-2969 b, TOI-2989 b, and TOI-5300 b. There are also new parameters for seven planets: KELT-11 b, Kepler-10 b, c, & d, MASCARA-4 b, and TOI-1468 b, WASP-127 b. We also have seven more spectra for five planets: GJ 357 b, HAT-P-2 b, LHS 1140 c, LTT 1445 A b, and WASP-127 b. You can find these in the .

An ExoFOP Milestone: 1 Million Files and Counting!

The Exoplanet Follow-up Observing Program (ExoFOP), which was created in 2008 to help researchers collaborate and share observational data to confirm exoplanet candidates, has surpassed one million user-submitted data files. This incredible achievement is a testament to the dedication and collaboration of the exoplanet research community. Of the nearly 6,000 exoplanets confirmed thus far, data shared by the community through ExoFOP has contributed to approximately 4,000 of them.

From all of us at NExScI, thank you for your continued support and commitment to exoplanet science. This milestone would not have been possible without you. Read more about this exciting milestone in IPAC's media release.

3 EXOPLANET ARCHIVES

June 5, 2025

It's Getting Spectra-ordinary Around Here

The archive has six new planets this week and 14 spectra, which inches us closer to another data milestone of 1,000 atmospheric spectra!

This week's update bumps up the Atmospheric Spectroscopy Table's spectra count to **979**, with new data for TOI-1468 b (4 spectra), WASP-43 b (1), TrES-4 b (2), WASP-15 b (4), and HD 219666 b (3).

Pro Tip: You can access the table from the planet's Planetary Parameters section in its System Overview page.

This week's new planets are TIC 434398831 b & c, KMT-2017-BLG-2197L b, KMT-2022-BLG-1790L b, KMT-2022-BLG-2076L b, and KMT-2023-BLG-2209L b.

We've also updated the disposition of 42 Dra b to False Positive Planet based on a published refutation. The object has been removed from the Planetary Systems and Planetary Systems Composite Parameters tables, but its data are still available on the System Overview page.

Our Accepted Paper is Now on arXiv!

We're happy to announce our latest paper, "The NASA Exoplanet Archive and Exoplanet Follow-up Observing Program: Data, Tools, and Usage," has been accepted by the *Planetary Science Journal* and is available in arXiv as Christiansen et al. (2025). This paper is an updated overview of the archive's data, services, and tools since Akeson et al. (2013), our initial publication.

Please use Christiansen et al. (2025) for all future citations. We will update the link to the paper on our Acknowledging the NASA Exoplanet Archive in Publications page once the paper is posted to ADS.

A Quick Update About Our Planetary Systems Composite Data Table

We've made a few changes to the Planetary Systems Composite Data (PSCompPars) table's underlying algorithms, which should improve the precision and accuracy of the displayed values. Users may notice new values for some parameters.

Download/Website: https://exoplanetarchive.ipac.caltech.edu *Contact:* mharbut@caltech.edu

4 Jobs and Positions

Courtois Scientific Vanguard Fund UdeM FAS Postdoctoral Fellowship at the Trottier Institute for Research on Exoplanets (IREx)

Prof. René Doyon

Montréal, Canada, Starting date: January to September 2026

The Trottier Institute for Research on Exoplanets (IREx) at the Université de Montréal invites applications for the **Courtois Scientific Vanguard Fund of the Faculty of Arts and Sciences Postdoctoral Competition**. Candidates in experimental, observational, or theoretical astrophysics related to exoplanets or astronomical instrumentation are encouraged to apply.

Six (6) fellowships are available in the natural and formal sciences (including, but not limited to, astrophysics), for **3 years**, offering a salary of up to CAD\$76,000/year and research funds up to CAD\$30,000/year, subject to justification.

Deadline: September 23, 2025, 11:59 p.m. (ET)

Applicants must have obtained their Ph.D. within the last 5 years (6 with justification) and hold **Canadian** citizenship, permanent residency, or a valid (or pending) work permit.

The research proposal must be original and integrated with a Faculty of Arts and Science researcher in natural/formal sciences. A faculty member must confirm their willingness to supervise. For that matter, astrophysicists working in exoplanet science should contact IREx director René Doyon at irex-applications@umontreal.ca as soon as possible.

See full details online: https://exoplanetes.umontreal.ca/en/job/courtois/.

IREx (https://exoplanetes.umontreal.ca/en/) is a dynamic team of over 60 researchers across Quebec, Canada (UdeM, McGill, Bishop's, Université Laval, Montreal Planetarium) working on cutting-edge observational, theoretical, and instrumental projects related to exoplanets. The team is deeply involved in major international efforts, including the *James Webb Space Telescope*, SPIRou, and NIRPS, with privileged access to data from these instruments.

IREx also leads a strong science communication and outreach program, training researchers who excel both scientifically and as communicators.

We value diversity, equity, and inclusion, and strongly encourage applications from underrepresented groups in physics. Our EDI committee supports the integration of these individuals into our research environment.

Download/Website: https://exoplanetes.umontreal.ca/en/job/courtois/\
Contact: irex-applications@umontreal.ca

4 JOBS AND POSITIONS

Postdoctoral Position on Exoplanet Demographics and Planet Formation

Gijs D. Mulders

Institute of Astrophysics, Pontificia Universidad Catolica, Santiago, Chile

Santiago, Chile, 2025 or early 2026

Applications are invited for a two-year postdoctoral position on the demographics and formation of exoplanets. The selected candidate will work in the group of Prof. Gijs Mulders at Pontificia Universidad Catolica (PUC) in Santiago, Chile on a project funded by the European Southern Observatory and the Government of Chile. The main focus of the project are the interactions between super-earths and giant planets, and is flexible to a more theoretical approach (planet formation simulations) or a more observational one (exoplanet demographics).

The Institute of Astrophysics at PUC is one of the leading astronomy centers in Latin America with 15 faculty, 20 postdocs, and 60 graduate students. While a resident at IA-PUC, the candidate has access to the Geryon 2 and 3 supercomputing faculties, and will be eligible for the 10% observing time on all telescopes in Chile. Postdocs at IA have the opportunity to teach undergraduate astronomy courses and mentor students.

The position is for 2 years, with a flexible starting date in 2025 or early 2026. Salary is 29.500.000 CLP/yr, commensurate with similar positions in the USA or Europe after cost of living adjustments. A budget for (international) travel, computing, and publications is available.

Applications should be uploaded by August 15th at https://form.jotform.com/251134649610049 and include:

- Cover letter
- CV
- Research statement describing past/current research
- Names of 2 or 3 references

Inquiries are welcome!

Download/Website: https://form.jotform.com/251134649610049 Contact: gijs.mulders@uc.cl

5 As seen on astro-ph

The following list contains exoplanet related entries appearing on astro-ph in June 2025.

Disclaimer: The hyperlinks to the astro-ph articles are provided for the convenience of the reader, but the ExoPlanet News cannot be responsible for their accuracy and perpetuity.

June 2025

astro-ph/2506.00091: AGNI: A radiative-convective model for lava planet atmospheres by Harrison Nicholls et al.

astro-ph/2506.00470: Star-Planet Interactions: A Computational View by A. A. Vidotto

- astro-ph/2506.00669: On Linking Planet Formation Models, Protoplanetary Disk Properties, and Mature Gas Giant Exoplanet Atmospheres *by Adina D. Feinstein et al.*
- astro-ph/2506.01132: NLTE atmospheric modelling of the ultra-hot Jupiter WASP-178b and comparison with UV and optical observations by L. Fossati et al.
- astro-ph/2506.02199: WASP-121b's transmission spectrum observed with JWST/NIRSpec G395H reveals thermal dissociation and SiO in the atmosphere by Cyril Gapp et al.
- astro-ph/2506.02188: Interior redox state effects on the stability of secondary atmospheres and observational manifestations: LP 791-18 d as a case study for outgassing rocky exoplanets by Leonardos Gkouvelis et al.
- astro-ph/2506.02144: A Pan-STARRS Search for Distant Planets: Part 1 by Matthew J. Holman et al.
- astro-ph/2506.02253: Stability of a cluster-disrupted mean-motion resonance (chain) in HR 8799 and PDS 70 by Brent Maas et al.
- astro-ph/2506.01800: A Precise Metallicity and Carbon-to-Oxygen Ratio for a Warm Giant Exoplanet from its Panchromatic JWST Emission Spectrum *by Lindsey S. Wiser et al.*
- astro-ph/2506.01771: SiO and a super-stellar C/O ratio in the atmosphere of the giant exoplanet WASP-121b by Thomas M. Evans-Soma et al.
- astro-ph/2506.02934: Eighteen Exoplanet Host Stars from the NPOI Data Archive by Ellyn K. Baines et al.
- astro-ph/2506.05392: Exoplaneteers Keep Overestimating Sigma Significances by David Kipping, Björn Benneke
- astro-ph/2506.03299: The NASA Exoplanet Archive and Exoplanet Follow-up Observing Program: Data, Tools, and Usage by Jessie L. Christiansen et al.
- astro-ph/2506.03248: New Rotation Periods from the Kepler Bonus Background Light Curves by Zachary R. Claytor, Jamie Tayar
- astro-ph/2506.02886: Stellar Magnetic Storm Induced Magnetospheric Polarity Reversals: Distinguishing between Unmagnetised and Magnetised Exoplanets by Sakshi Gupta et al.
- astro-ph/2506.02826: Hot exozodiacal dust around Fomalhaut: The MATISSE perspective by Kevin Ollmann et al.
- astro-ph/2506.02748: MINDS: The very low-mass star and brown dwarf sample. Detections and trends in the inner disk gas by A. M. Arabhavi et al.
- astro-ph/2506.02721: Clues for Solar System Formation from Meteorites and their Parent Bodies by Bernard Marty et al.
- astro-ph/2506.04415: Modelling shadows in scattered light observations as signals from companions in protoplanetary discs by Deniz Akansoy et al.
- astro-ph/2506.04348: Planet Earth in reflected and polarized light: II. Refining contrast estimates for rocky exoplanets with ELT and HWO by Giulia Roccetti et al.
- astro-ph/2506.04199: Abundance Effects from Protoplanetary Disk Outflows by Åke Nordlund
- astro-ph/2506.04440: The Age and High Energy Environment of the Very Young Transiting Exoplanet TOI 1227b by Attila Varga et al.

- astro-ph/2506.03624: Peering through the veil: Investigating protoplanetary disk outer edges using backside visibility by Joel George et al.
- astro-ph/2506.03511: POLARIS: A High-contrast Polarimetric Imaging Benchmark Dataset for Exoplanetary Disk Representation Learning by Fangyi Cao et al.
- astro-ph/2506.04164: A Resonant Beginning for the Solar System Terrestrial Planets by Shuo Huang et al.
- astro-ph/2506.05631: The TESS Ten Thousand Catalog: 10,001 uniformly-vetted and -validated Eclipsing Binary Stars detected in Full-Frame Image data by machine learning and analyzed by citizen scientists
 - by Veselin B. Kostov et al.
- astro-ph/2506.05521: The TOI-1117 Multi-planetary System: 3 sub-Neptunes, 1 in both the Neptunian Desert and Radius Valley by Isobel S. Lockley et al.
- astro-ph/2506.05556: **DART-Vetter: A Deep LeARning Tool for automatic triage of exoplanet candidates** by Stefano Fiscale et al.
- astro-ph/2506.04923: Three Hot Jupiters transiting K-dwarfs with a significant heavy element mass by Y. G. C. Frensch et al.
- astro-ph/2506.04604: **Radio prospects of extrasolar aurorae polaris as a probe of planetary magnetism** *by Asaf Kaya, Tansu Daylan*
- astro-ph/2506.05089: Planets similar in size are often dissimilar in interior by E. Mamonova et al.
- astro-ph/2506.04605: Searching for helium escape and a low density atmosphere around the 120 Myr old sub-Neptune HIP94235b using CRIRES+ by Ava Morrissey et al.
- astro-ph/2506.05892: Disk Evolution Study Through Imaging of Nearby Young Stars (DESTINYS): Evidence of planet-disk interaction in the 2MASSJ16120668-3010270 system by C. Ginski et al.
- astro-ph/2506.06195: TOI-2407 b: a warm Neptune in the desert by C. Janó Muñoz et al.
- astro-ph/2506.06429: Maximizing Ariel's Survey Leverage for Population-Level Studies of Exoplanets by Nicolas B. Cowan, Ben Coull-Neveu
- astro-ph/2506.06433: Grand Design Spiral Arms in the Compact, Embedded Protoplanetary Disk of Haro 6-13 by Jane Huang et al.
- astro-ph/2506.06445: Multi-band, Multi-epoch Photometry of the Spot-crossing System TOI-3884: Refined System Geometry and Spot Properties by Mayuko Mori et al.
- astro-ph/2506.06865: FAUST XXVI. The dust opacity spectral indices of protostellar envelopes bridge the gap between interstellar medium and disks by Luca Cacciapuoti et al.
- astro-ph/2506.07265: Limb-Darkening Coefficients for the 4-Term and Power-2 Laws for the JWST Space Mission, Adopting PHOENIX Spherical Models at High Resolution *by A. Claret et al.*
- astro-ph/2506.08232: A Multi-Species Atmospheric Escape Model with Excited Hydrogen and Helium: Application to HD209458b by Anna Ruth Taylor et al.
- astro-ph/2506.08177: Constraining Ongoing Volcanic Outgassing Rates and Interior Compositions of Extrasolar Planets with Mass Measurements of Plasma Tori *by V. Abby Boehm et al.*
- astro-ph/2506.08088: Planetesimal Scattering Efficiency of Cold Giant Planet Architectures by Stephen R. Kane, Emma L. Miles
- astro-ph/2506.07898: Constraints on Quaoar's rings and atmosphere from JWST/NIRCam observations of a stellar occultation by Benjamin Proudfoot et al.
- astro-ph/2506.08014: Why M-dwarf flares have limited impact on sub-Neptunes' atmospheric evaporation by *Andrea Caldiroli et al.*
- astro-ph/2506.07912: A circularly polarized low-frequency radio burst from the exoplanetary system HD 189733 by X. Zhang et al.
- astro-ph/2506.08078: The Epoch of Giant Planet Migration Planet Search Program. III. The Occurrence Rate of Young Giant Planets Inside the Water Ice Line by Quang H. Tran et al.
- astro-ph/2506.07931: A transiting giant planet in orbit around a 0.2-solar-mass host star by Edward M. Bryant et al.
- astro-ph/2506.08499: Impact chronology of leftover planetesimals by R. Brasser

- astro-ph/2506.08798: Characterization of the Visual Binary TOI-6883AB and its dynamical implications for the planetary companion TOI-6883Ab *by G. Conzo et al.*
- astro-ph/2506.08814: SPECULOOS: five years hunting terrestrial planets around ultra-cool dwarfs by Sebastián Zúñiga-Fernández et al.
- astro-ph/2506.08969: A New Spectral Class of Brown Dwarfs at the Bottom of the IMF in IC 348 by K. L. Luhman, C. Alves de Oliveira
- astro-ph/2506.09201: JWST Coronagraphic Images of 14 Her c: a Cold Giant Planet in a Dynamically Hot, Multi-planet System by Daniella C. Bardalez Gagliuffi et al.
- astro-ph/2506.09976: MINDS: Detection of an inner gas disk caused by evaporating bodies around HD 172555 *by M. Samland et al.*
- astro-ph/2506.11200: DBNets2.0: simulation-based inference for planet-induced dust substructures in protoplanetary discs by A. Ruzza et al.
- astro-ph/2506.10969: The Gap-Giant Association: Are Planets Hiding in the Gaps? by Caleb Lammers, Joshua N. Winn
- astro-ph/2506.10812: Self-gravity in thin protoplanetary discs: 1. The smoothing-length approximation versus the exact self-gravity kernel by S. Rendon Restrepo et al.
- astro-ph/2506.10752: The ALMA Survey of Gas Evolution of PROtoplanetary Disks (AGE-PRO): XII. Extreme millimetre variability detected in a Class II disc by James M. Miley et al.
- astro-ph/2506.10750: The ALMA Survey of Gas Evolution of PROtoplanetary Disks (AGE-PRO): XI. Beamcorrected gas disk sizes from fitting 12CO moment zero maps by Leon Trapman et al.
- astro-ph/2506.10746: The ALMA Survey of Gas Evolution of PROtoplanetary Disks (AGE-PRO): X. Dust Substructures, Disk Geometries, and Dust-disk Radii by Miguel Vioque et al.
- astro-ph/2506.10743: The ALMA Survey of Gas Evolution of PROtoplanetary Disks (AGE-PRO): VIII. The impact of external photoevaporation on disk masses and radii in Upper Scorpius by Rossella Anania et al.
- astro-ph/2506.10742: The ALMA Survey of Gas Evolution of PROtoplanetary Disks (AGE-PRO): VII. Testing accretion mechanisms from disk population synthesis by Benoît Tabone et al.
- astro-ph/2506.10738: The ALMA Survey of Gas Evolution of PROtoplanetary Disks (AGE-PRO): V. Protoplanetary gas disk masses by Leon Trapman et al.
- astro-ph/2506.10740: The ALMA Survey of Gas Evolution of PROtoplanetary Disks (AGE-PRO): VI. Comparison of Dust Evolution Models to AGE-PRO Observations *by Nicolas T. Kurtovic et al.*
- astro-ph/2506.10435: On the Mass Budget Problem of Protoplanetary Disks: Streaming Instability and Optically Thick Emission by Daniel Godines et al.
- astro-ph/2506.10735: The ALMA Survey of Gas Evolution of PROtoplanetary Disks (AGE-PRO): IV. Dust and Gas Disk Properties in the Upper Scorpius Star-forming Region by Carolina Agurto-Gangas et al.
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