# Domain 2

## Planetary systems architecture, their formation and evolution



Swiss National Science Foundation



## PlanetS AG, 28 April 2022



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Coordinator: X. Dumusque

Signal correction to reveal other Earths X. Dumusque (UGE) **ERC** funded



Understanding planetary system architectures S. Udry (UGE)

2.1





2.4

Direct imaging of forming and mature planetary systems S. Quanz (ETHZ)



Planet evolution: connection with stellar properties, processes, and formation R. Helled (UZH)







Monitoring solar system formation and evolution with noble gases H. Busemann (ETHZ)



Collisional growth and orbital evolution of planets J. Stadel (UZH)



Formation & composition of core dominated planets Y. Alibert (UBE)

**Domain 2 projects** 

**Externally funded** 











**Domain 2 projects** 

**Externally funded** 







## **Planet**S

## 8 Domain 2 generic meetings 4 Domain 2 specific workshops



**Externally funded** 









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	P 2.1	P 2.2	P 2.3	P 2.4	P 2.5	P 2.6	P 2.7	P 2.8
WP1	In progress	Nearly Completed	Completed	In progress				
WP2	In progress	Nearly Completed	In progress					
WP3	In progress	In progress		In progress	In progress	In progress	Completed	In progress
WP4	Refocused	In progress			Completed	In progress	In progress	In progress
WP5	Not started				In progress			Soon starting

FNSNF Swiss National Science Foundation

## **Planet**S



## Publications



Swiss National Science Foundation

## Planet S'

## 31 papers with first author payed by NCCR (~1 paper/year/project) 81 papers with major contribution from NCCR (~2.5 papers/year/project)



## Highlights since last site visit

## Understanding planetary system architectures

## Planet S





### Understanding planetary system architectures

## Before optimising TTV detection

![](_page_12_Figure_2.jpeg)

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### After

![](_page_12_Figure_6.jpeg)

![](_page_12_Picture_7.jpeg)

![](_page_13_Picture_1.jpeg)

![](_page_13_Picture_2.jpeg)

![](_page_13_Picture_3.jpeg)

![](_page_13_Picture_5.jpeg)

NIRPS integrated in Chile

Commissioning starts June 7th

Data flow, Data reduction, configuration for simultaneous observations with HARPS completed

![](_page_14_Picture_4.jpeg)

![](_page_14_Picture_5.jpeg)

## Direct imaging of forming and mature planetary systems

![](_page_15_Picture_1.jpeg)

![](_page_15_Picture_2.jpeg)

![](_page_15_Picture_3.jpeg)

Atmospheric retrieval code that use cross-correlation and combines photometry, low and medium resolution spectroscopy observation of directly imaged exoplanets

CRODODILE Hayoz et al. 2022, in prep.

![](_page_16_Picture_3.jpeg)

![](_page_16_Figure_5.jpeg)

![](_page_16_Picture_8.jpeg)

![](_page_17_Picture_1.jpeg)

![](_page_17_Picture_2.jpeg)

![](_page_17_Picture_3.jpeg)

![](_page_17_Picture_5.jpeg)

## Venus internal structure

![](_page_18_Figure_2.jpeg)

![](_page_18_Picture_4.jpeg)

## Core size of Venus between 2930 and 4350 km for all assumed composition

## Similarity of multiplanet systems Otegi et al. 21

![](_page_19_Figure_2.jpeg)

Planets in systems that are similar in radius could be different in mass and vice-versa More similarities in radius than in mass

A transition in the "peas in the pod" pattern for planets with M > ~100 M $_{\oplus}$  and R > ~10 R $_{\oplus}$ .

![](_page_19_Picture_5.jpeg)

## Planet S

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![](_page_20_Picture_4.jpeg)

## Planet S

![](_page_21_Picture_4.jpeg)

First populations based on hybrid (pebbles/planetesimal) formation model

## Planet S

![](_page_22_Picture_6.jpeg)

First populations based on hybrid (pebbles/planetesimal) formation model

First models with planetesimal fragmentation and drift computed

![](_page_23_Picture_3.jpeg)

![](_page_23_Picture_7.jpeg)

First populations based on hybrid (pebbles/planetesimal) formation model

First models with planetesimal fragmentation and drift computed

Bayesian internal structure models applied to numerous CHEOPS and ESPRESSO planets

![](_page_24_Picture_4.jpeg)

![](_page_24_Picture_8.jpeg)

First populations based on hybrid (pebbles/planetesimal) formation model

First models with planetesimal fragmentation and drift computed

Bayesian internal structure models applied to numerous CHEOPS and ESPRESSO planets

Model of circumstellar disk (Yuhito Shibaike's presentation)

![](_page_25_Picture_5.jpeg)

![](_page_25_Picture_8.jpeg)

![](_page_26_Figure_1.jpeg)

Intra-system correlations used to identify Earth-harboring systems Work by Jean Davoult (based on results from Mishra et al. 22, sub.)

![](_page_26_Picture_3.jpeg)

![](_page_26_Picture_4.jpeg)

## Collisional growth and orbital evolution of planets

![](_page_27_Picture_1.jpeg)

![](_page_27_Picture_3.jpeg)

![](_page_27_Picture_4.jpeg)

## **Collisional growth and orbital evolution of planets**

## New GENGA 2.0 code released (Grimm et al. 22, subm.)

Thomas Meier

![](_page_28_Picture_3.jpeg)

### **PI: Joachim Stadel**

![](_page_28_Picture_7.jpeg)

## **Collisional growth and orbital evolution of planets**

Thomas Meier

Smooth particle hydrodynamics added to state-of-the-art PKDGRAV3 code Many applications, like high-resolution model to study the formation of the Moon

Planet S

![](_page_29_Picture_4.jpeg)

![](_page_29_Picture_5.jpeg)

![](_page_29_Picture_6.jpeg)

Planet S

### PI: Henner Busemann

![](_page_30_Picture_3.jpeg)

## Back to the lab

![](_page_31_Picture_2.jpeg)

### PI: Henner Busemann

![](_page_31_Picture_4.jpeg)

Back to the lab

Planet S

Measurements of Ryugu dust samples returned with JAXA's Hayabusa 2 spacecraft (Science, in review)

![](_page_32_Figure_4.jpeg)

![](_page_32_Picture_6.jpeg)

Back to the lab

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Measurements of Ryugu dust samples returned with JAXA's Hayabusa 2 spacecraft (Science, in review)

Measurements of gas from the Hayabusa 2 spacecraft, outgassed from recovered samples (Science Adv., in review)

### **PI: Henner Busemann**

![](_page_33_Picture_6.jpeg)

## **Stellar Correction to Reaveal other Earths**

![](_page_34_Picture_1.jpeg)

![](_page_34_Picture_2.jpeg)

![](_page_34_Picture_3.jpeg)

![](_page_34_Picture_4.jpeg)

## **Stellar Correction to Reaveal other Earths**

![](_page_35_Figure_1.jpeg)

Modelisation of stellar activity in high-resolution spectra

Allows to investigate stellar activity correction techniques at the spectral level

![](_page_35_Picture_4.jpeg)

![](_page_35_Picture_5.jpeg)

![](_page_35_Picture_6.jpeg)

![](_page_36_Picture_0.jpeg)